

## The thermal-statistical model for particle production

Jean Cleymans<sup>1,a</sup>

UCT-CERN Research Centre and Department of Physics,  
University of Cape Town, South Africa

**Abstract.** We introduce the thermal statistical model for particle production and its applications to heavy-ion collision experiments. We focus on the aspects related to chemical equilibrium, in particular on particle yields.

### Acknowledgment

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<sup>a</sup> e-mail: Jean.Cleymans@uct.ac.za

# The Thermal-Statistical Model for Particle Production I.

J. Cleymans

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## Outline

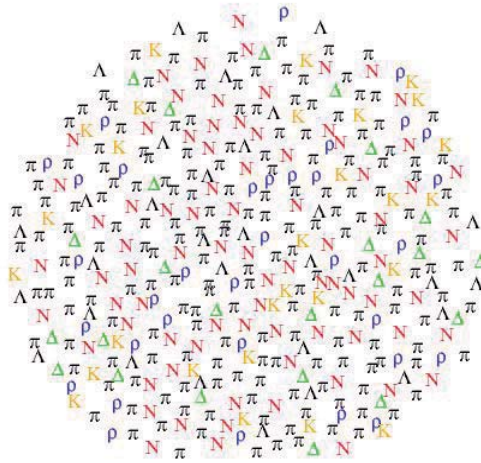
Statistical Model

Strangeness

$E_T/N_{ch}$  vs.  $E/N$

The Horn in the  $K^+/\pi^+$  Ratio





J.C. and H. Satz, Zeitschrift fuer Physik C57, 135 (1993)



Statistical Model      Strangeness       $E_T / N_{ch}$  vs.  $E/N$       The Horn in the  $K^+ / \pi^+$  Ratio

## Thermal Equilibrium

In thermal equilibrium

$$Z = \text{Tr} e^{-\frac{H}{T} + \frac{\mu N}{T}}$$

$$\langle N \rangle = \frac{\text{Tr} N e^{-\frac{H}{T} + \frac{\mu N}{T}}}{\text{Tr} e^{-\frac{H}{T} + \frac{\mu N}{T}}}$$

$$\langle E \rangle = \frac{\text{Tr} E e^{-\frac{H}{T} + \frac{\mu N}{T}}}{\text{Tr} e^{-\frac{H}{T} + \frac{\mu N}{T}}}$$



## Full Hydrodynamic Flow

### Bjorken scaling + Transverse expansion

After integration over  $m_T$

$$\frac{dN_i/dy}{dN_j/dy} = \frac{N_i^0}{N_j^0}$$

where  $N_i^0$  is the particle yield  
as calculated in a fireball **AT REST!**

**Effects of hydrodynamic flow cancel out in ratios.**



## Thermal Equilibrium

Particle Number

$$\begin{aligned} \langle N \rangle &= \frac{\text{Tr} N e^{-\frac{H}{T} + \frac{\mu N}{T}}}{\text{Tr} e^{-\frac{H}{T} + \frac{\mu N}{T}}} \\ &= \frac{T}{Z} \frac{\partial}{\partial \mu} \text{Tr} e^{-\frac{H}{T} + \frac{\mu N}{T}} \\ &= T \frac{1}{Z} \frac{\partial Z}{\partial \mu} \\ &= T \frac{\partial}{\partial \mu} \ln Z \end{aligned}$$



## Thermal Equilibrium

### Average Energy

$$\begin{aligned}
 \langle E \rangle &= \frac{\text{Tr } H e^{-\frac{H}{T} + \frac{\mu N}{T}}}{\text{Tr } e^{-\frac{H}{T} + \frac{\mu N}{T}}} \\
 &= -\frac{1}{Z} \frac{\partial Z}{\partial \beta} + \mu \langle N \rangle \\
 &= T^2 \frac{\partial}{\partial T} \ln Z + \mu \langle N \rangle
 \end{aligned}$$



## Thermal Equilibrium

$$\begin{aligned}
 N_i &= g_i V \int \frac{d^3 p}{(2\pi)^3} \exp\left(-\frac{E}{T}\right) e^{\frac{\mu_i}{T}} \\
 &= g_i V \frac{4\pi}{(2\pi)^3} \int p^2 dp \exp\left(-\frac{\sqrt{p^2 + m_i^2}}{T}\right) e^{\frac{\mu_i}{T}} \\
 &= g_i V \frac{4\pi}{(2\pi)^3} T^3 \int x^2 dx \exp\left(-\sqrt{x^2 + m_i^2/T^2}\right) e^{\frac{\mu_i}{T}} \\
 &= g_i V \frac{1}{2\pi^2} T m_i^2 K_2\left(\frac{m_i}{T}\right) e^{\frac{\mu_i}{T}}
 \end{aligned}$$



## Thermal Equilibrium

$$n_i = g_i \frac{1}{2\pi^2} T m_i^2 K_2\left(\frac{m_i}{T}\right) e^{\frac{\mu_i}{T}}$$

$$\epsilon_i = g_i \frac{1}{2\pi^2} T m_i^3 \left[ K_1\left(\frac{m_i}{T}\right) + 3 \frac{T}{m} K_2\left(\frac{m_i}{T}\right) \right] e^{\frac{\mu_i}{T}}$$

$$s_i = g_i \frac{1}{2\pi^2} m_i^3 \left[ K_1\left(\frac{m_i}{T}\right) + \frac{4T}{m} K_2\left(\frac{m_i}{T}\right) - \frac{\mu_i}{m} K_2\left(\frac{m_i}{T}\right) \right] e^{\frac{\mu_i}{T}}$$

$$P_i = g_i \frac{1}{2\pi^2} T^2 m_i^2 K_2\left(\frac{m_i}{T}\right) e^{\frac{\mu_i}{T}}$$



## Chemical Equilibrium

In equilibrium

$$E_1 + E_2 + \dots = E_3 + E_4 + E_5 + \dots \quad (1)$$

for the chemical potentials

$$\mu_1 + \mu_2 + \dots = \mu_3 + \mu_4 + \mu_5 + \dots \quad (2)$$

As an example

$$\pi^0 + p \leftrightarrow \pi_0 + p + \pi^0 \quad (3)$$

leads to

$$\mu_{\pi^0} + \mu_p = \mu_{\pi^0} + \mu_p + \mu_{\pi^0} \quad (4)$$

which leads to

$$\mu_{\pi^0} = 0 \quad (5)$$



## Chemical Equilibrium

In equilibrium

$$B + \bar{B} \leftrightarrow B + B + B + \bar{B} \quad (6)$$

$$dE = -pdV + TdS + \mu_B dN_B + \mu_{\bar{B}} dN_{\bar{B}}$$

Due to baryon number conservation one has

$$N_B - N_{\bar{B}} = \text{constant}$$

and

$$dN_B = dN_{\bar{B}}$$

The energy is a minimum for

$$dE = (\mu_B + \mu_{\bar{B}}) dN_B = 0 \quad (7)$$

$$\mu_B = -\mu_{\bar{B}} \quad (8)$$



## Chemical Equilibrium

In equilibrium

$$N_B = g V \int \frac{d^3p}{(2\pi)^3} \exp\left(-\frac{E}{T} + \frac{\mu_B}{T}\right)$$

$$N_{\bar{B}} = g V \int \frac{d^3p}{(2\pi)^3} \exp\left(-\frac{E}{T} - \frac{\mu_B}{T}\right)$$

$$N_B = N_{\bar{B}} \rightarrow \mu_B = 0$$

$$N_B \geq N_{\bar{B}} \rightarrow \mu_B \geq 0$$

$$N_B \leq N_{\bar{B}} \rightarrow \mu_B \leq 0$$



	Chemical Equilibrium	No Chem. Equil.
$\pi$	$\exp\left[-\frac{E_\pi}{T}\right]$	$\exp\left[-\frac{E_\pi}{T} + \frac{\mu_\pi}{T}\right]$
$N$	$\exp\left[-\frac{E_N}{T} + \frac{\mu_B}{T}\right]$	$\exp\left[-\frac{E_N}{T} + \frac{\mu_N}{T}\right]$
$\bar{N}$	$\exp\left[-\frac{E_N}{T} - \frac{\mu_B}{T}\right]$	$\exp\left[-\frac{E_N}{T} + \frac{\mu_{\bar{N}}}{T}\right]$
$\Lambda$	$\exp\left[-\frac{E_\Lambda}{T} + \frac{\mu_B}{T} - \frac{\mu_S}{T}\right]$	$\exp\left[-\frac{E_\Lambda}{T} + \frac{\mu_\Lambda}{T}\right]$
$\bar{\Lambda}$	$\exp\left[-\frac{E_\Lambda}{T} - \frac{\mu_B}{T} + \frac{\mu_S}{T}\right]$	$\exp\left[-\frac{E_\Lambda}{T} + \frac{\mu_{\bar{\Lambda}}}{T}\right]$
$K$	$\exp\left[-\frac{E_K}{T} + \frac{\mu_S}{T}\right]$	$\exp\left[-\frac{E_K}{T} + \frac{\mu_K}{T}\right]$
$\bar{K}$	$\exp\left[-\frac{E_K}{T} - \frac{\mu_S}{T}\right]$	$\exp\left[-\frac{E_K}{T} + \frac{\mu_{\bar{K}}}{T}\right]$



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The number of particles of type  $i$  is determined by:

$$N_i = V g_i \int \frac{d^3 p}{(2\pi)^3} \exp\left(-\frac{E_i}{T} + \frac{\mu_i}{T}\right)$$

For bosons:

$$N_i = V g_i \int \frac{d^3 p}{(2\pi)^3} \frac{1}{\exp\left(\frac{E_i}{T} - \frac{\mu_i}{T}\right) - 1}$$

For fermions:

$$N_i = V g_i \int \frac{d^3 p}{(2\pi)^3} \frac{1}{\exp\left(\frac{E_i}{T} - \frac{\mu_i}{T}\right) + 1}$$

Only conserved quantum numbers matter for chemical equilibrium: In equilibrium

$$\mu_i = B_i \mu_B + Q_i \mu_Q + S_i \mu_S + C_i \mu_C + \dots \quad (9)$$



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Statistical Model      Strangeness       $E_T / N_{ch}$  vs.  $E/N$       The Horn in the  $K^+ / \pi^+$  Ratio

$g_i$	$m_i$	stat	$S_i$	$B_i$	$Q_i$	Particle $i$
1	0.140	-1	0	0	1.	$\pi^+$
1	0.135	-1	0	0	0.	$\pi^0$
1	0.140	-1	0	0	-1.	$\pi^-$
1	0.547	-1	0	0	0.	$\eta$
3	0.770	-1	0	0	1.	$\rho^+$
3	0.770	-1	0	0	0.	$\rho^0$
3	0.770	-1	0	0	-1.	$\rho^-$
3	0.782	-1	0	0	0.	$\omega$
1	0.958	-1	0	0	0.	$\eta'$
1	0.980	-1	0	0	0.	$f_0$
1	0.982	-1	0	0	1.	$a_0^+$
1	0.982	-1	0	0	0.	$a_0^0$
1	0.982	-1	0	0	-1.	$a_0^-$
3	1.019	-1	0	0	0.	$\phi$
3	1.170	-1	0	0	0.	
3	1.230	-1	0	0	1.	
3	1.230	-1	0	0	0.	
3	1.230	-1	0	0	-1.	
3	1.229	-1	0	0	1.	
3	1.229	-1	0	0	0.	
3	1.229	-1	0	0	-1.	
5	1.275	-1	0	0	0.	
3	1.282	-1	0	0	0.	
1	1.297	-1	0	0	0.	
1	1.300	-1	0	0	1.	
1	1.300	-1	0	0	0.	

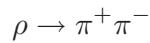


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Statistical Model      Strangeness       $E_T / N_{ch}$  vs.  $E/N$       The Horn in the  $K^+ / \pi^+$  Ratio

## The Role of Resonances

### Example: $\rho$ 's



Final, observed, number of  $\pi^+$  is given by

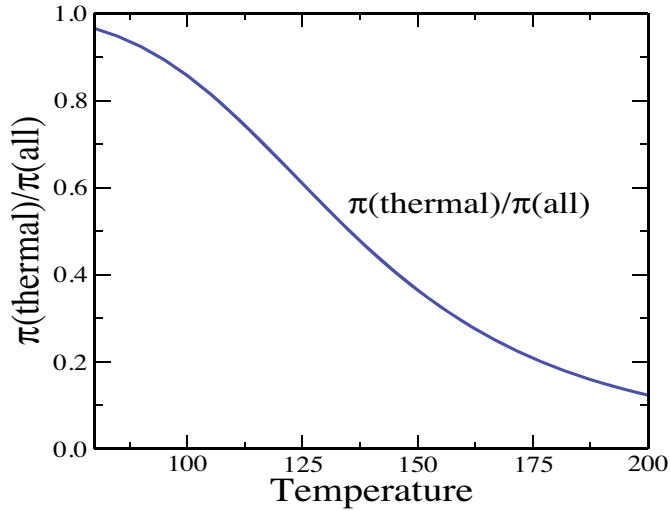
$$N_{\pi^+} = N_{\pi^+}(\text{thermal}) + N_{\pi^+}(\text{resonance decays})$$

depending on the temperature, over 80% of observed pions are due to resonance decays

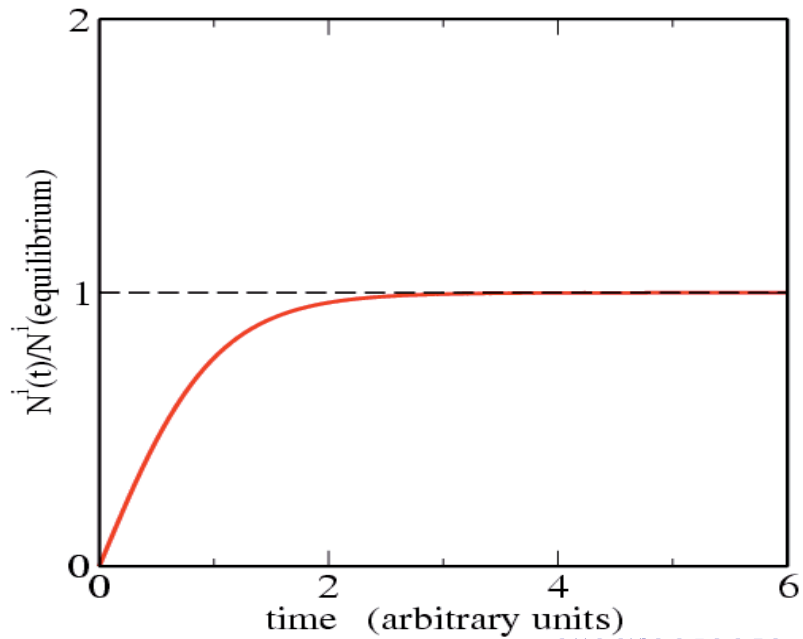


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## Importance of Resonances.



## Strangeness saturation?



## Strangeness saturation?

$$N_i = \boxed{\gamma_s^{|S_i|}} V g_i \int \frac{d^3p}{(2\pi)^3} \exp\left(-\frac{E_i}{T} + \frac{\mu_i}{T}\right)$$

with

$\gamma_s < 1$  strangeness under-saturation

$\gamma_s = 1$  strangeness in chemical equilibrium

$\gamma_s > 1$  strangeness over-saturation



## SPS data.

	Measurement
Pb-Pb 158A GeV	
$(\pi^+ + \pi^-)/2.$	$600 \pm 30$
$K^+$	$95 \pm 10$
$K^-$	$50 \pm 5$
$K_S^0$	$60 \pm 12$
$p$	$140 \pm 12$
$\bar{p}$	$10 \pm 1.7$
$\phi$	$7.6 \pm 1.1$
$\Xi^-$	$4.42 \pm 0.31$
$\Xi^-$	$0.74 \pm 0.04$
$\bar{\Lambda}/\Lambda$	$0.2 \pm 0.04$



## SPS data.

SPS: Chemical Freeze-Out Parameters:

$$\begin{aligned}
 T &= 156.0 \pm 2.4 \text{ MeV} \\
 \mu_B &= 239 \pm 12 \text{ MeV} \\
 \gamma_s &= 0.862 \pm 0.036
 \end{aligned}$$

F. Becattini, J.C., A. Keränen, E. Suhonen and K. Redlich  
 Physical Review C64 (2001) 024901.



## AGS data.

	Measurement
Au–Au 11.6A GeV	
Participants	$363 \pm 10$
$K^+$	$23.7 \pm 2.9$
$K^-$	$3.76 \pm 0.47$
$\pi^+$	$133.7 \pm 9.9$
$\Lambda$	$20.34 \pm 2.74$
$p/\pi^+$	$1.234 \pm 0.126$
$\bar{p}$	$>0.0185 \pm 0.0018$



## AGS data.

AGS: Chemical Freeze-Out Parameters:

$$\begin{aligned}
 T &= 130.6 \pm 5.5 \text{ MeV} \\
 \mu_B &= 594 \pm 26 \text{ MeV} \\
 \gamma_s &= 0.883 \pm 0.124
 \end{aligned}$$

F. Becattini, J.C., A. Keränen, E. Suhonen and K. Redlich  
 Physical Review C64 (2001) 024901.



## SIS data.

	Measurement
Au–Au 1.7A GeV	
$\pi^+/\text{p}$	$0.052 \pm 0.013$
$K^+/\pi^+$	$0.003 \pm 0.00075$
$\pi^-/\pi^+$	$2.05 \pm 0.51$
$\eta/\pi^0$	$0.018 \pm 0.007$

SIS: Chemical Freeze-Out Parameters:

$$\begin{aligned}
 T &= 49.7 \pm 1.1 \text{ MeV} \\
 \mu_B &= 818 \pm 15 \text{ MeV} \\
 \gamma_s &= 1 \text{ (fixed)}
 \end{aligned}$$

J. C., H. Oeschler and K. Redlich  
 Physical Review C59, (1999) 1663.



## RHIC data.

J. C., B. Kämpfer, M. Kaneta, S. Wheaton, N. Xu, Phys. Rev. C71, 0409071 (2005)

Ratio	Experiment	Central	Mid-Central	Peripheral
$\pi_{(2)}^-/\pi_{(2)}^+$	BRAHMS	0.990±0.100		
	PHENIX	0.960±0.177	0.920±0.170	0.933±0.172
	PHOBOS	1.000±0.022		
	STAR	1.000±0.073	1.000±0.073	1.000 ± 0.073
$K_{(2)}^+/K_{(2)}^-$	PHENIX	1.152±0.240	1.292±0.268	1.322±0.284
	PHOBOS	1.099±0.111		
	STAR	1.109±0.022	1.105±0.036	1.120±0.040
$\bar{p}_{(1)}/p_{(1)}$	PHENIX	0.680±0.149	0.671±0.142	0.717±0.157
$\bar{p}_{(2)}/p_{(2)}$	BRAHMS	0.650±0.092		
	PHOBOS	0.600±0.072		
	STAR	0.714±0.050	0.724±0.050	0.764±0.053
$\bar{\Lambda}_{(1)}/\Lambda_{(1)}$	PHENIX	0.750±0.180	0.798±0.197	0.795±0.197
$\bar{\Lambda}_{(2)}/\Lambda_{(2)}$	STAR	0.719±0.090	0.739±0.092	0.744±0.100
$\Xi_{(2)}^+/\Xi_{(2)}^-$	STAR	0.840±0.053	0.822±0.114	0.815±0.096
$\bar{\Omega}^+/\Omega^-$	STAR	1.062±0.410		
$K_{(2)}^-/\pi_{(2)}^-$	PHENIX	0.151±0.030	0.134±0.027	0.116±0.023
	STAR	0.151±0.022	0.147±0.022	0.130±0.019
$K_S^0/\pi_{(2)}^-$	STAR	0.134±0.022	0.131±0.022	0.108±0.018
$\bar{p}_{(1)}/\pi_{(2)}^-$	PHENIX	0.049±0.010	0.047±0.010	0.045±0.009
$\bar{p}_{(2)}/\pi_{(2)}^-$	STAR	0.069±0.019	0.067±0.019	0.067±0.019
$\Lambda_{(1)}/\pi_{(2)}^-$	STAR	0.043±0.008	0.043±0.008	0.039±0.007
$\Lambda_{(2)}/\pi_{(2)}^-$	PHENIX	0.072±0.017	0.068±0.016	0.074±0.017
$\langle K^{*0} \rangle / \pi_{(2)}^-$	STAR	0.039±0.011		
$\phi/\pi_{(2)}^-$	STAR	0.022±0.003	0.021±0.004	0.022±0.004
$\Xi_{(2)}^-/\pi_{(2)}^-$	STAR	0.0093±0.0012	0.0072±0.0011	0.0060±0.0008



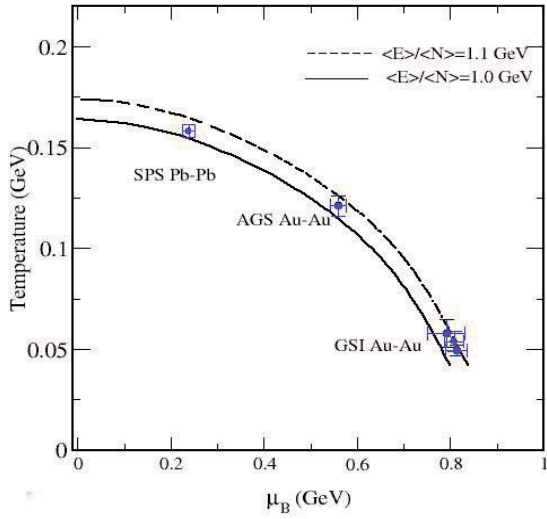
## RHIC data.

RHIC: Chemical Freeze-Out Parameters:

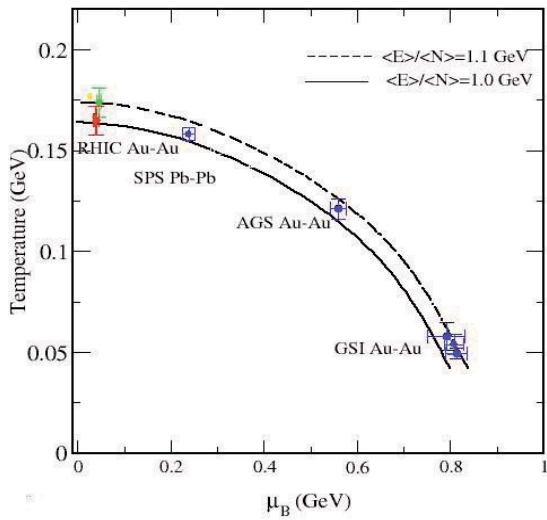
$$\begin{aligned}
 T &= 169 \pm 4.2 \text{ MeV} \\
 \mu_B &= 39.6 \pm 6 \text{ MeV} \\
 \gamma_s &= 0.9 \pm 0.1
 \end{aligned}$$

J. C., B. Kämpfer, M. Kaneta, S. Wheaton, N. Xu  
 Phys. Rev. C71, 0409071 (2005)



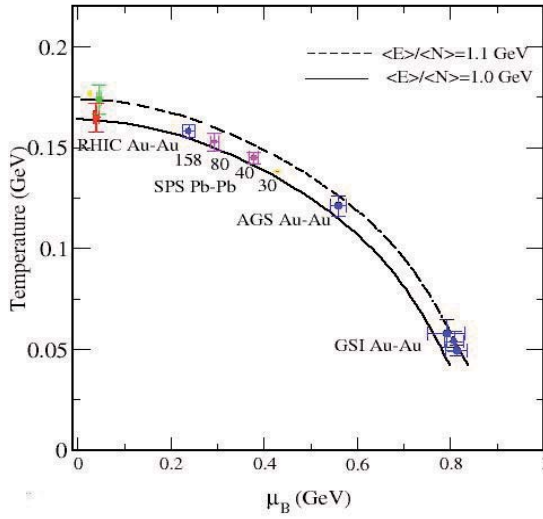


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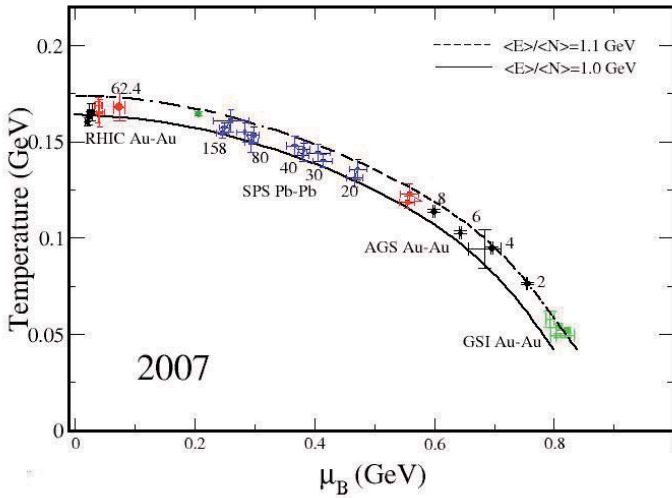


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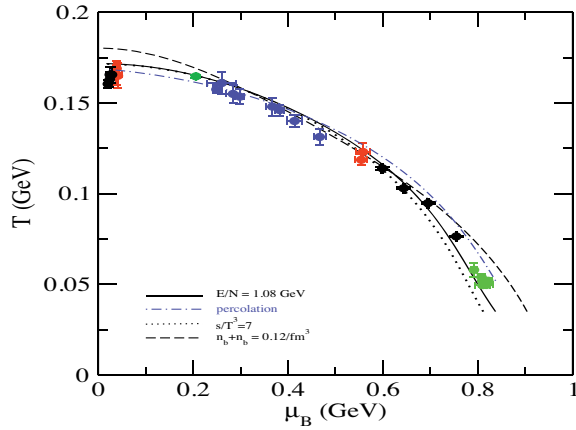
# E/N in 2005



# E/N in 2007

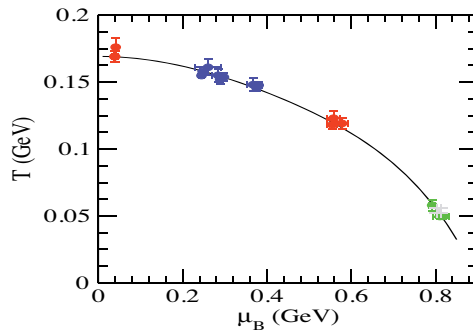






V. Magas and H. Satz, Eur. Phys. J. **C32** 115 (2003).

P. Braun-Munzinger and J. Stachel, J. Phys. G: Nucl. Part. Phys. **28** 1971 (2002).

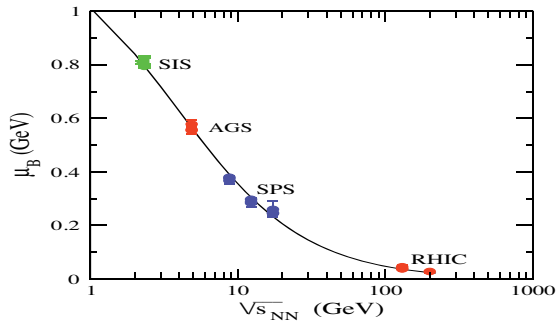


$$T(\mu_B) = 0.169 - 0.189\mu_B^2 + 0.165\mu_B^4 - 0.229\mu_B^6.$$

J. C., H. Oeschler, K. Redlich, S. Wheaton



$\mu_B$  as a function of  $\sqrt{s_{NN}}$



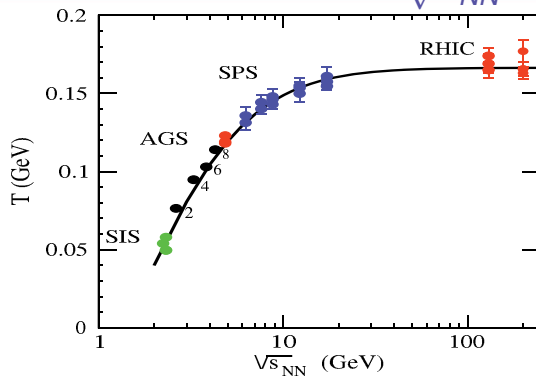
$$\mu_B(\sqrt{s}) = \frac{1.273 \text{ GeV}}{1 + 0.258 \text{ GeV}^{-1} \sqrt{s}}$$

This predicts at LHC  $\mu_B \approx 1 \text{ MeV}$ .

J. C., H. Oeschler, K. Redlich, S. Wheaton



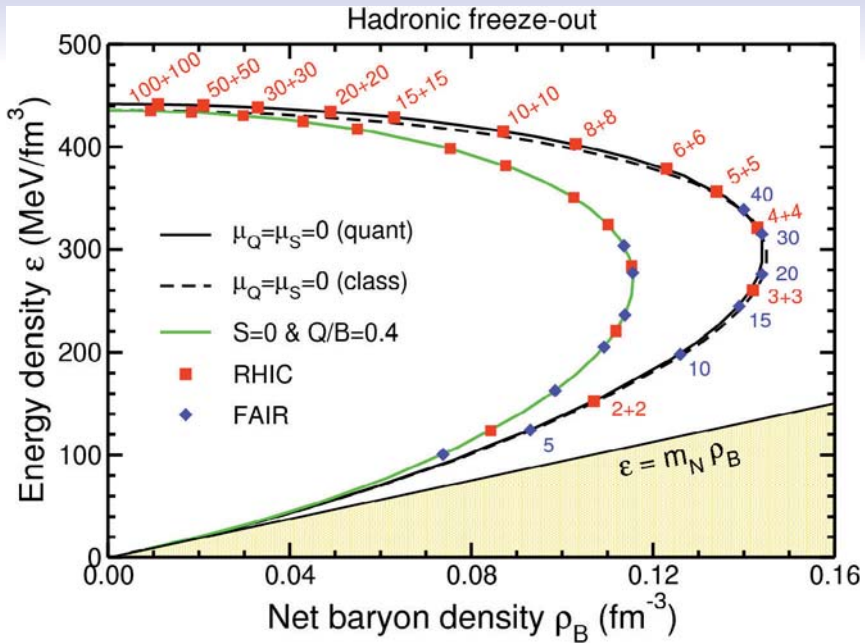
$T$  as a function of  $\sqrt{s_{NN}}$



This predicts at LHC  $T \approx 170 \text{ MeV}$ .

J. C., H. Oeschler, K. Redlich, S. Wheaton





J. Randrup and J.C., Phys. Rev. C74 (2006) 047901

### Will it be possible to determine directly $E/N$ ?

$E$ : energy of primordial hadrons

$N$ : number of primordial hadrons

$$\begin{aligned} \langle E_T \rangle &= \langle E \sin \theta \rangle \\ &= \frac{\pi}{4} \langle E \rangle \end{aligned}$$

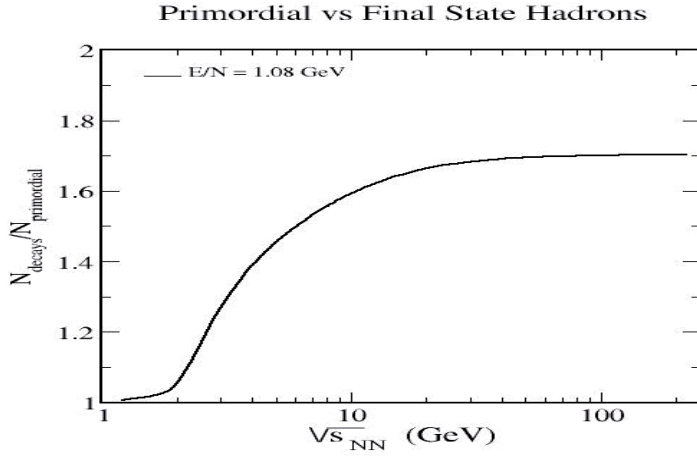
Low energy limit

$$\lim \frac{E_T}{N_{ch}} = \frac{\frac{\pi}{4} m_N}{0.4} \approx 1.8 \text{ GeV}$$

High energy limit

$$\lim \frac{E_T}{N_{ch}} = \frac{\frac{\pi}{4} \langle M \rangle}{2/3} \approx 0.9 \text{ GeV}$$

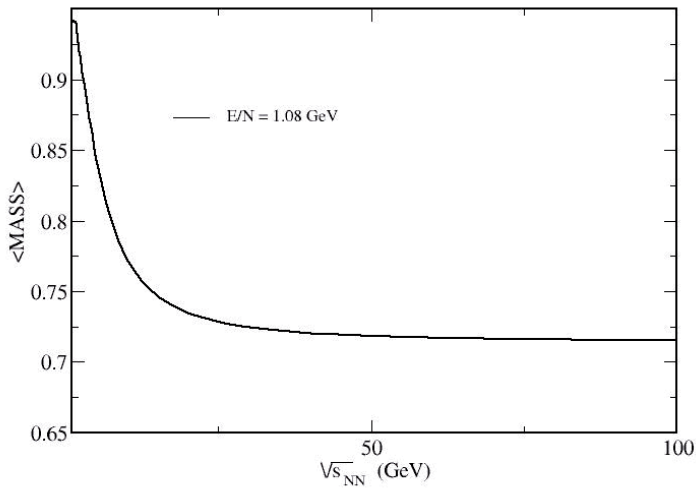
However  $E_T$  : subtract  $m_N$  for baryons add  $m_N$  for antibaryons.



J.C., R. Sahoo, D.K. Srivastava, S. Wheaton



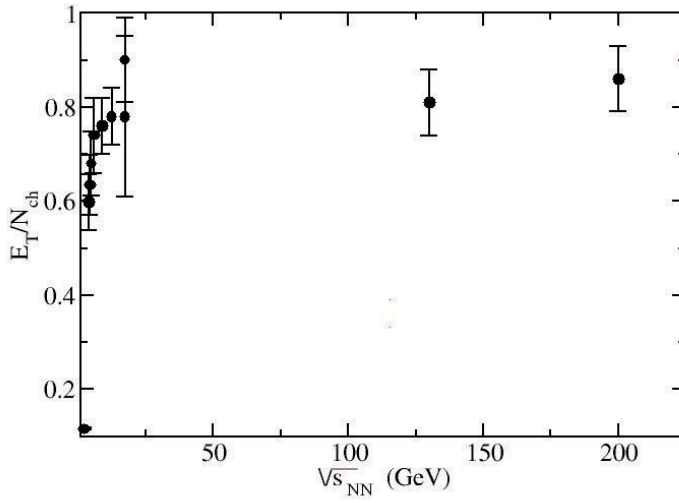
### Average Mass in Fireball



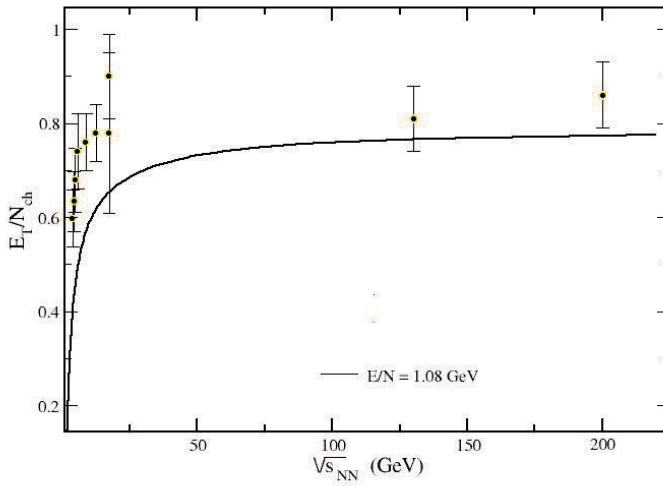
J.C., R. Sahoo, D.K. Srivastava, S. Wheaton



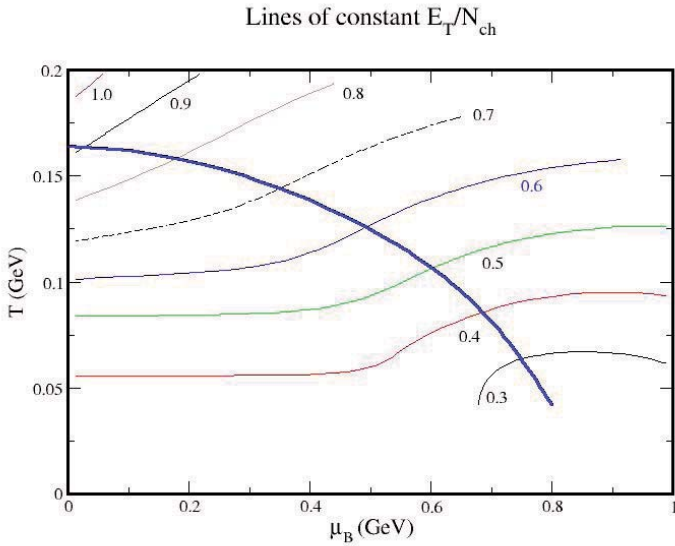
Transverse Energy per Charge



Transverse Energy per Charged Hadron

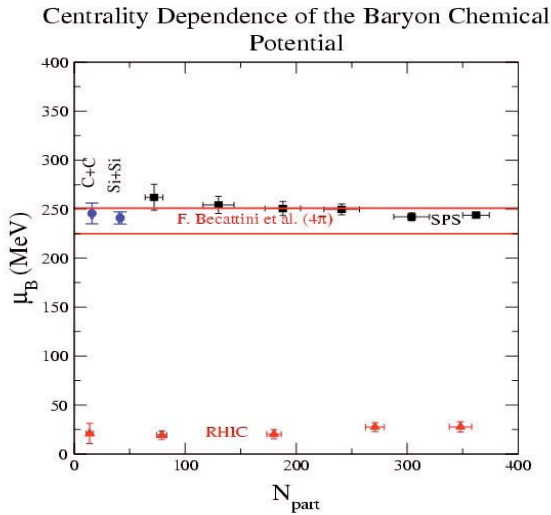


J.C., R. Sahoo, D.K. Srivastava, S. Wheaton



mainly follows  $T$  and is determined by  $E/N$ ,  
 J.C., R. Sahoo, D.K. Srivastava, S. Wheaton

$E_T/N_{ch}$



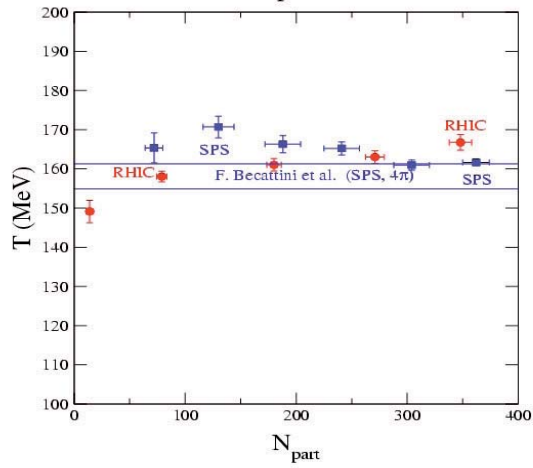
J. C., B. Kämpfer, P. Steinberg and S. Wheaton, Journal of Physics G30 S595-S598 (2004).



Statistical Model

$K^+/\pi^+$  Ratio

Centrality Dependence of the Chemical Freeze-out Temperature



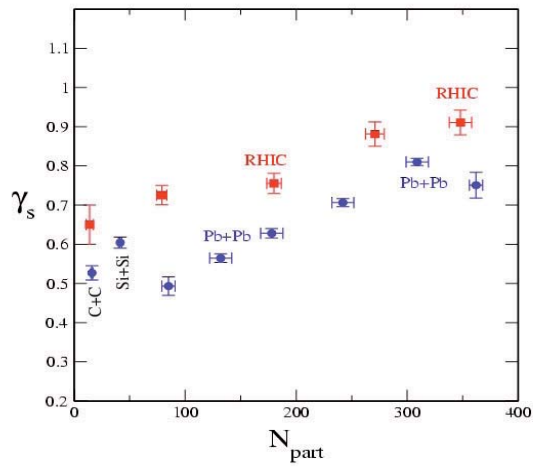
J. C., B. Kämpfer, P. Steinberg and S. Wheaton, Journal of Physics G30 S595-S598 (2004).



Navigation icons: back, forward, search, etc.

Statistical Model

$K^+/\pi^+$  Ratio



J. C., B. Kämpfer, P. Steinberg and S. Wheaton, Journal of Physics G30 S595-S598 (2004).



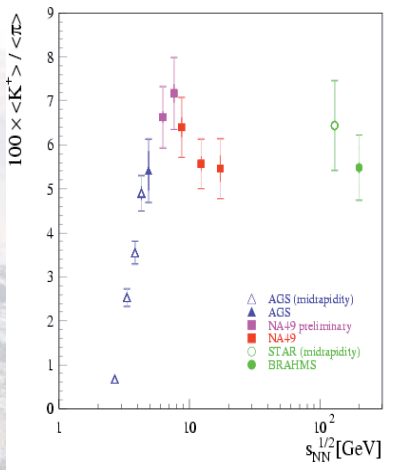
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The NA49 Collaboration has recently performed a series of measurements of Pb-Pb collisions at 20, 30, 40, 80 and 158 AGeV beam energies . When these results are combined with measurements at lower beam energies from the AGS they reveal an unusually sharp variation with beam energy in the  $\Lambda / \langle \pi \rangle$ , with  $\langle \pi \rangle \equiv 3/2(\pi^+ + \pi^-)$ , and  $K^+ / \pi^+$  ratios. Such a strong variation with energy does not occur in pp collisions and therefore indicates a major difference in heavy-ion collisions. This transition has been referred as the “horn”.



# The Elephant in the Room

Friese  
Dinkelaker  
Blume  
Speltz



Difficult to avoid, Hard to Model  
 → But no unambiguous corroborating evidence





Strangeness in Heavy Ion Collisions  
vs  
Strangeness in pp - collisions

Use the Wroblewski factor

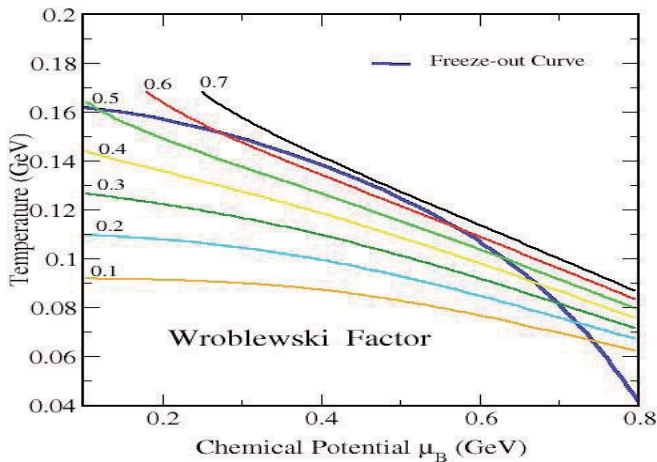
$$\lambda_s = \frac{2 \langle s\bar{s} \rangle}{\langle u\bar{u} \rangle + \langle d\bar{d} \rangle}$$

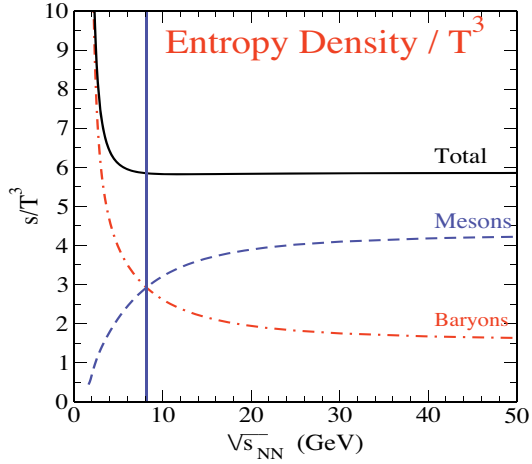
This is determined by the number of **newly** created quark – anti-quark pairs and **before** strong decays, i.e. before  $\rho$ 's and  $\Delta$ 's decay.

Limiting values :

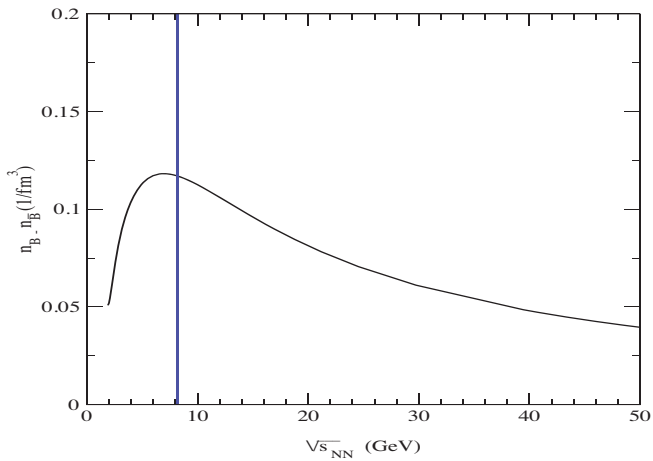
$\lambda_s = 1$  all quark pairs are equally abundant, SU(3) symmetry.

$\lambda_s = 0$  no strange quark pairs.





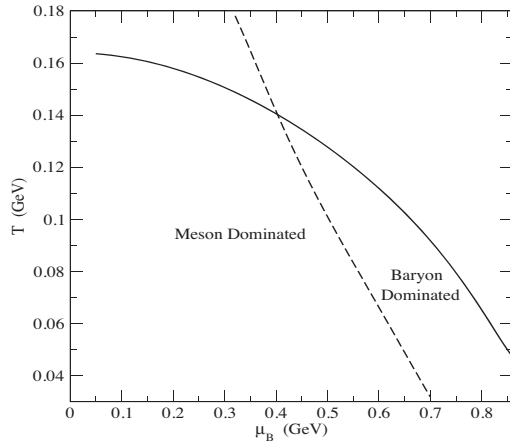
J. C., H. Oeschler, K. Redlich and S. Wheaton, Physics Letters B615 (2005) 50-54.



J. C., H. Oeschler, K. Redlich and S. Wheaton, Physics Letters B615 (2005) 50-54.



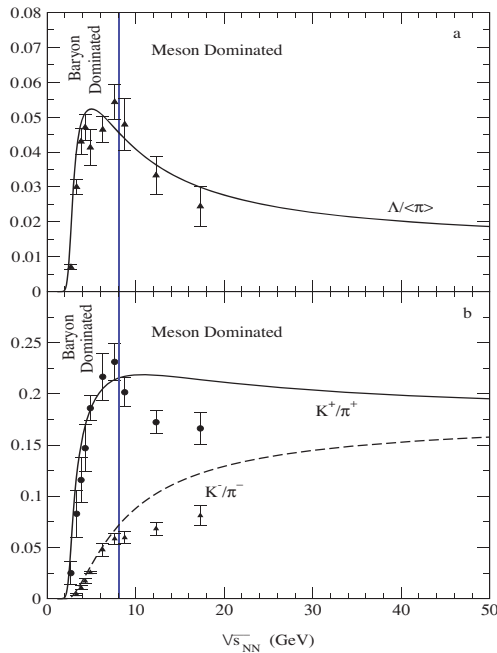
# Dense Matter In Heavy Ion Collisions and Astrophysics (DM2008)



J. C., H. Oeschler, K. Redlich and S. Wheaton, Physics Letters B615 (2005) 50-54.



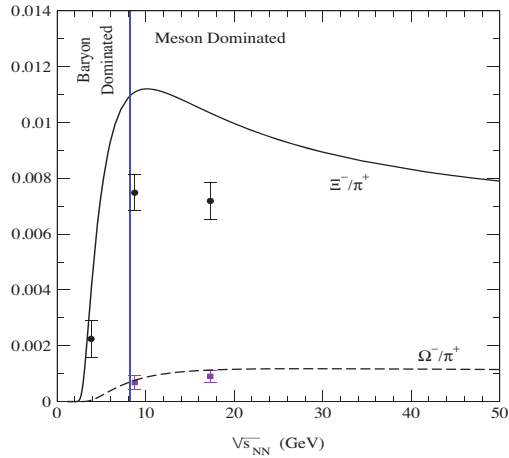
Navigation icons: back, forward, search, etc.



J. C., H. Oeschler, K. Redlich and S. Wheaton, Physics Letters B615 (2005) 50-54.



Navigation icons: back, forward, search, etc.



J. C., H. Oeschler, K. Redlich and S. Wheaton, Physics Letters B615 (2005) 50-54.



Maxima in Particle Ratios predicted by the Thermal Model.

Ratio	Maximum at $\sqrt{s_{NN}}$ (GeV)	Maximum Value
$\Lambda/\langle\pi\rangle$	5.1	0.052
$\Xi^-/\pi^+$	10.2	0.011
$K^+/\pi^+$	10.8	0.22
$\Omega^-/\pi^+$	27	0.0012

J. C., H. Oeschler, K. Redlich and S. Wheaton, Physics Letters B615 (2005) 50-54.



# The Thermal-Statistical Model for Particle Production II.

J. Cleymans

25 July 2008 / Dubna, Russia



## Outline

Hydrodynamic Flow

Excluded Volume Corrections

Canonical Corrections

Transverse Momentum Distributions

Rapidity Distributions



## Hydrodynamic Flow.

### Cooper-Frye formula

From

$$N = \int j^\mu d\sigma_\mu \quad \text{and} \quad j^\mu = \int d^3p \frac{p^\mu}{E} f(r, p, t)$$

obtain

$$E \frac{dN}{d^3p} = \int p^\mu d\sigma_\mu f(r, p, t)$$



The number of particles of type  $i$  is determined by:

$$E \frac{dN_i}{d^3p} = \frac{g_i}{(2\pi)^3} \int d\sigma_\mu p^\mu \exp\left(-\frac{p^\mu u_\mu}{T} + \frac{\mu_i}{T}\right)$$

Integrating this over all momenta

$$N_i = \frac{g_i}{(2\pi)^3} \int d\sigma_\mu \int \frac{d^3p}{E} p^\mu \exp\left(-\frac{p^\mu u_\mu}{T} + \frac{\mu_i}{T}\right)$$

or

$$N_i = \int d\sigma_\mu u^\mu n_i(T, \mu)$$

If the temperature and chemical potential are unique along the freeze-out curve

### Particle Yield:

$$N_i = n_i(T, \mu) \int d\sigma_\mu u^\mu$$

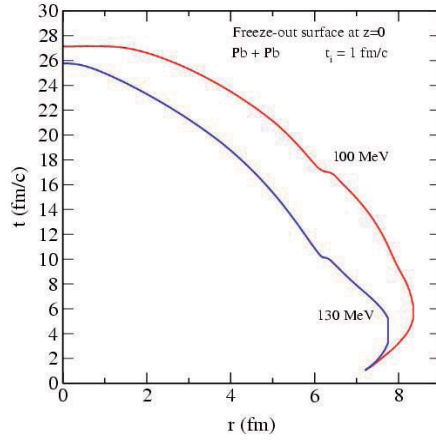
i.e. integrated  $(4\pi)$  multiplicities are the same as for a single fireball at rest (apart from the volume).



# Dense Matter In Heavy Ion Collisions and Astrophysics (DM2008)

Hydrodynamic Flow Excluded Volume Corrections

Transverse Momentum Distributions Rapidity Distributions

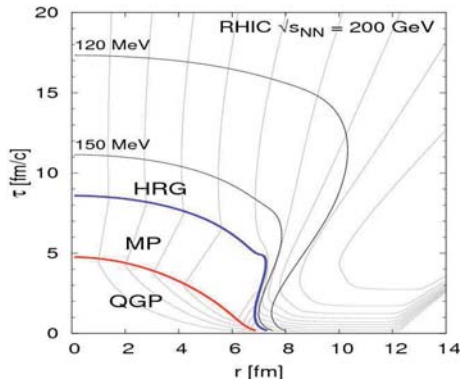


J. Cleymans, K. Redlich, D.K. Srivastava  
Phys. Rev. C55 (1997) 1431



Navigation icons: back, forward, search, etc.

Hydrodynamic Flow Excluded Volume Corrections Canonical Corrections Transverse Momentum Distributions Rapidity Distributions



K. Eskola, H. Honkanen, H.Niemi, P.V. Ruuskanen, S.S. Räsänen, hep-ph/0506049



Navigation icons: back, forward, search, etc.

## Full Hydrodynamic Flow

Bjorken scaling + Transverse expansion after integration over  $m_T$

$$\left(\frac{dN_i}{dy}\right)_{y=0} = \frac{g}{\pi} \int_{\sigma} r dr \tau_F(r) \left\{ \cosh(y_T) - \left(\frac{\partial \tau_F}{\partial r}\right) \sinh(y_T) \right\} m_i^2 TK_2\left(\frac{m_i}{T}\right)$$

Consequence :  $\frac{dN_i/dy}{dN_j/dy} = \frac{N_i^0}{N_j^0}$

**Effects of hydrodynamic flow cancel out in ratio.**



## Excluded Volume Corrections.

$$\begin{aligned} Z &= \exp \left\{ V \int \frac{d^3 p}{(2\pi)^3} e^{-\frac{\epsilon}{T} + \frac{\mu}{T}} \right\} \\ &= \sum_{N=0}^{\infty} \frac{V^N}{N!} e^{\mu N/T} \left[ \int \frac{d^3 p}{(2\pi)^3} e^{-\frac{\epsilon}{T}} \right]^N \end{aligned}$$

with excluded volume corrections

$$\begin{aligned} Z &\rightarrow \sum_{N=0}^{\infty} \frac{(V - V_0 N)^N}{N!} e^{\mu N/T} \\ &\quad \left[ \int \frac{d^3 p}{(2\pi)^3} e^{-\frac{\epsilon}{T}} \right]^N \theta(V - V_0 N) \end{aligned}$$





## Excluded Volume Corrections.

It is more convenient to consider these corrections in the pressure ensemble:

$$Z_p \equiv \int_0^\infty dV e^{-PV/T} \sum_{N=0}^\infty \frac{V^N}{N!} e^{\mu N/T} \left[ \int \frac{d^3p}{(2\pi)^3} e^{-\frac{\epsilon}{T}} \right]^N$$

$$\begin{aligned} Z_p &\rightarrow \sum_{N=0}^\infty \int_0^\infty dV e^{-PV/T} \\ &\quad \frac{(V - V_0 N)^N}{N!} e^{\mu N/T} \\ &\quad \left[ \int \frac{d^3p}{(2\pi)^3} e^{-\frac{\epsilon}{T}} \right]^N \theta(V - V_0 N) \end{aligned}$$

introduce  $x \equiv V - V_0 N$ .



## Excluded Volume Corrections.

$$\begin{aligned} Z_p &= \sum_{N=0}^\infty \int_0^\infty dx e^{-Px/T} \\ &\quad \frac{x^N}{N!} e^{-PV_0 N/T} e^{\mu N/T} \\ &\quad \left[ \int \frac{d^3p}{(2\pi)^3} e^{-\frac{\epsilon}{T}} \right]^N \end{aligned}$$

a new variable  $\bar{\mu} \equiv \mu - PV_0$



## Excluded Volume Corrections.

$$Z_p \rightarrow \sum_{N=0}^{\infty} \int_0^{\infty} dx e^{-Px/T} \frac{x^N}{N!} e^{\bar{\mu}N/T} \left[ \int \frac{d^3p}{(2\pi)^3} e^{-\frac{\epsilon}{T}} \right]^N$$

which is the original partition function with the replacement

$$\bar{\mu} = \mu - P V_0$$



## Excluded Volume Corrections.

The particle number density now becomes:

$$\begin{aligned} n &= \frac{T}{V} \frac{\partial}{\partial \mu} \ln Z \\ &= \frac{T}{V} \frac{\partial \bar{\mu}}{\partial \mu} \frac{\partial}{\partial \bar{\mu}} \ln Z \\ &= \frac{\partial \bar{\mu}}{\partial \mu} n_0 \\ &= [1 - V_0 n] n_0 \end{aligned}$$

$$n = \frac{n_0}{1 + V_0 n_0}$$

Effects Cancel Out in Ratios.

J.C., K. Redlich, H. Satz, E. Suhonen, ZfP C33, 151, (1986)

D.H. Rischke, M.I. Gorenstein, H. Stöcker, W. Greiner, ZfP C51, 485 (1991).



## Exact Strangeness Conservation.

For a small system at low temperatures ( $T \approx 50$  MeV), e.g. at GSI canonical corrections are necessary.

Instead of

$$N_K \approx \exp -M_K/T$$

one gets

$$N_K \approx \exp -2M_K/T$$

Extra suppression is due to strangeness conservation and lack of a large heat bath. This correction disappears quickly at higher energies and is already small at AGS energies.



## Exact Strangeness Conservation.

$$Z = \text{Tr} e^{-\frac{H}{T} + \frac{\mu N}{T}}$$

Insert a Kronecker delta in the trace:

$$\begin{aligned} & \sum_i n_i(S=1) + 2 \sum_j n_j(S=2) + 3 \sum_k n_k(S=3) = \\ & \sum_i \bar{n}_i(S=-1) + 2 \sum_j \bar{n}_j(S=-2) + 3 \sum_k \bar{n}_k(S=-3) \end{aligned}$$

and rewrite it as

$$\begin{aligned} & \delta \left( \sum_i n_i(S=1) + \dots, \sum_i \bar{n}_i(S=-1) + \dots \right) \\ & = \frac{1}{2\pi} \int_0^{2\pi} d\phi \\ & \exp \left( i\phi \sum_i n_i(S=1) + \dots - i\phi \sum_i \bar{n}_i(S=-1) \right) \end{aligned}$$



## Exact Strangeness Conservation.

$$\begin{aligned}
 Z &= \frac{1}{2\pi} \int_0^{2\pi} d\phi \exp \left\{ Z_1 e^{i\phi} + Z_{-1} e^{-i\phi} \right. \\
 &\quad + Z_2 e^{2i\phi} + Z_{-2} e^{-2i\phi} \\
 &\quad \left. + Z_3 e^{3i\phi} + Z_{-3} e^{-3i\phi} \right\} \\
 &= \frac{1}{2\pi} \int_0^{2\pi} d\phi \exp \left\{ \sqrt{Z_1 Z_{-1}} \left[ \sqrt{\frac{Z_1}{Z_{-1}}} e^{i\phi} + \sqrt{\frac{Z_{-1}}{Z_1}} e^{-i\phi} \right] \right. \\
 &\quad + \sqrt{Z_2 Z_{-2}} \left[ \sqrt{\frac{Z_2}{Z_{-2}}} e^{2i\phi} + \sqrt{\frac{Z_{-2}}{Z_2}} e^{-2i\phi} \right] \\
 &\quad \left. + \sqrt{Z_3 Z_{-3}} \left[ \sqrt{\frac{Z_3}{Z_{-3}}} e^{3i\phi} + \sqrt{\frac{Z_{-3}}{Z_3}} e^{-3i\phi} \right] \right\}
 \end{aligned}$$

$Z_i$ : sum of all particles with strangeness 1. e.a.  $K^+$

## Exact Strangeness Conservation.

Use

$$\exp \left\{ \frac{x}{2} \left( t + \frac{1}{t} \right) \right\} = \sum_{n=-\infty}^{\infty} I_n(x) t^n$$

to obtain

$$\begin{aligned}
 Z &= \frac{1}{2\pi} \int_0^{2\pi} e^{3im\phi + 2in\phi + ip\phi} \\
 &\quad \sum_{p=-\infty}^{\infty} I_p(x_1) \sum_{n=-\infty}^{\infty} I_n(x_2) \sum_{m=-\infty}^{\infty} I_m(x_3) \\
 &\quad y_1^p y_2^n y_3^m
 \end{aligned}$$

where

$$y_i = \sqrt{\frac{Z_i}{Z_{-i}}} \quad x_i = 2\sqrt{Z_i Z_{-i}}$$

## Exact Strangeness Conservation.

$$Z = \sum_{m=-\infty}^{\infty} \sum_{n=-\infty}^{\infty} I_{3m+2n}(x_1) y_1^{-3m-2n} I_n(x_2) y_2^n I_m(x_3) y_3^m$$



## Exact Strangeness Conservation.

$$\begin{aligned} Z &= \frac{1}{2\pi} \int_0^{2\pi} d\phi \exp \left\{ Z_1 e^{i\phi} + Z_{-1} e^{-i\phi} \right\} \\ &= \frac{1}{2\pi} \int_0^{2\pi} d\phi \exp \left\{ \sqrt{Z_1 Z_{-1}} \left[ \sqrt{\frac{Z_1}{Z_{-1}}} e^{i\phi} + \sqrt{\frac{Z_{-1}}{Z_1}} e^{-i\phi} \right] \right\} \end{aligned}$$

$Z_1$ : sum of all particles with strangeness 1, e.g.  $K^+$   
 $Z_{-1}$ : sum of all particles with strangeness -1, e.g.  $\Lambda$



## Exact Strangeness Conservation.

Use

$$\exp \left\{ \frac{x}{2} \left( t + \frac{1}{t} \right) \right\} = \sum_{n=-\infty}^{\infty} I_n(x) t^n$$

to obtain

$$Z = \frac{1}{2\pi} \int_0^{2\pi} e^{ip\phi} \sum_{p=-\infty}^{\infty} I_p(x_1) y_1^p$$

where

$$y_1 = \sqrt{\frac{Z_1}{Z_{-1}}} \quad x_1 = 2\sqrt{Z_1 Z_{-1}}$$

$$Z = I_0(x_1)$$



## Exact Strangeness Conservation.

In more detail, e.g. the multiplicity of  $K^+$

$$N_{K^+} = \frac{T}{Z} \frac{\partial I_0(x_1)}{\partial \mu_{K^+}} \Big|_{\mu_{K^+}=0}$$

Use

$$\frac{d}{dz} I_0(z) = I_1(z)$$



## Exact Strangeness Conservation.

$$\begin{aligned}
 N_{K^+} &= \frac{T}{Z} \frac{\partial}{\partial \mu_{K^+}} I_0(x_1) \\
 &= \frac{T}{I_0(x_1)} I_1(x_1) \frac{\partial x_1}{\partial \mu_{K^+}} \\
 &= \frac{T}{I_0(x_1)} I_1(x_1) \frac{\partial 2\sqrt{Z_1 Z_{-1}}}{\partial \mu_{K^+}} \\
 &= \frac{I_1(x_1)}{I_0(x_1)} \sqrt{\frac{Z_{-1}}{Z_1}} N_{K^+}^0
 \end{aligned}$$

where  $N_{K^+}^0$  refers to the "unmodified" kaon multiplicity.



## Exact Strangeness Conservation.

In the small volume limit this becomes

$$\lim_{z \rightarrow 0} I_0(z) = 1$$

and

$$\lim_{z \rightarrow 0} I_1(z) = \frac{Z}{2}$$

$$\begin{aligned}
 \lim_{V \rightarrow 0} &= N_{K^+}^0 Z_{-1} \\
 \lim &= N_{K^+}^0 Z_{-1} \\
 &= N_{K^+}^0 [N_{K^-}^0 + N_{\Lambda}^0 + \dots]
 \end{aligned}$$

i.e., the particle multiplicity is

- proportional to  $V^2$ , and not  $V^1$ .
- proportional to  $\exp(-2m_K/T)$  or to  $\exp(-(m_K + m_{\Lambda})/T)$  and not simply  $\exp(-m_K/T)$ , i.e. there is additional suppression of strange particles.



## Exact Strangeness Conservation.

In the small volume limit this becomes

$$\lim_{z \rightarrow 0} I_0(z) = 1$$

and

$$\lim_{z \rightarrow 0} I_1(z) = \frac{z}{2}$$

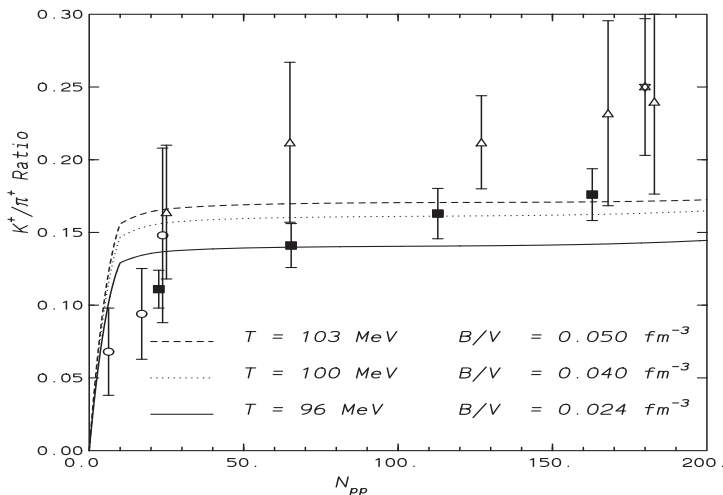
$$\begin{aligned} \lim_{V \rightarrow 0} &= N_{K^+}^0 Z_{-1} \\ \lim &= N_{K^+}^0 Z_{-1} \\ &= N_{K^+}^0 \left[ N_{K^-}^0 + N_{\Lambda}^0 + \dots \right] \end{aligned}$$

i.e., the particle multiplicity is

- proportional to  $V^2$ , and not  $V^1$ .
- proportional to  $\exp(-2m_K/T)$  or to  $\exp(-(m_K + m_{\Lambda})/T)$  and not simply  $\exp(-m_K/T)$ , i.e. there is additional suppression of strange particles.



## Exact Strangeness Conservation





## Transverse Momentum in the Thermal Model

$$E_i \frac{dN_i}{d^3p} = \frac{g_i}{(2\pi)^3} V E_i \exp\left(-\frac{E_i}{T} + \frac{\mu_i}{T}\right)$$

$$\frac{dN_i}{dy m_T dm_T} = \frac{g_i}{(2\pi)^2} V m_T \cosh y e^{-\frac{m_T}{T} \cosh y + \frac{\mu_i}{T}}$$

$$\frac{dN_i}{m_T dm_T} = \frac{g_i}{2\pi^2} V m_T K_1\left(\frac{m_T}{T}\right) e^{\frac{\mu_i}{T}}$$

For large values of  $m_T$  :

$$\frac{dN_i}{dm_T} = \frac{g_i}{2\pi^2} V m_T^{3/2} \exp\left(-\frac{m_T}{T}\right) \sqrt{\frac{T\pi}{2}} e^{\frac{\mu_i}{T}}$$

SCALING in  $m_T$



## Rapidity Distribution in the Thermal Model

$$\frac{dN_i}{dy m_T dm_T} = \frac{g_i}{(2\pi)^2} V m_T \cosh y e^{-\frac{m_T}{T} \cosh y + \frac{\mu_i}{T}}$$

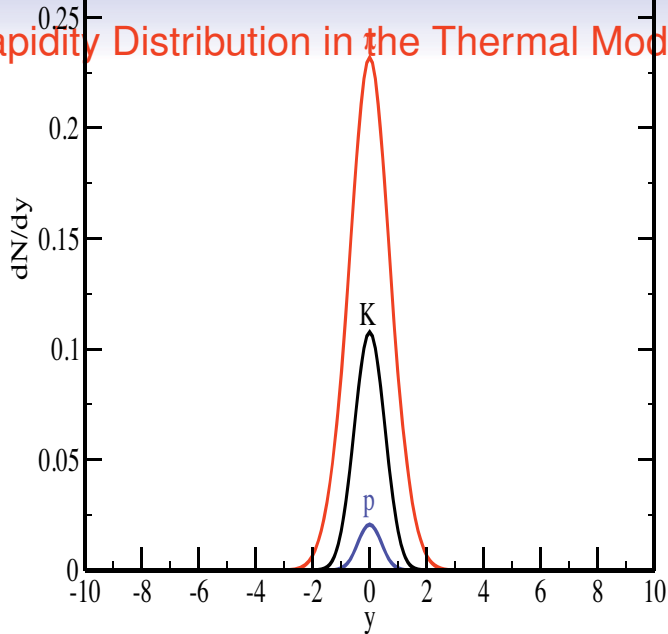
$$\frac{dN_i}{dy} = \frac{g_i V}{2\pi^2} \left[ \frac{2T^3}{\cosh^2 y} + \frac{2m_T^2}{\cosh y} + m^2 T \right] e^{\frac{\mu_i}{T}}$$

$$e^{-\frac{m}{T} \cosh y}$$

Narrow Distribution in Rapidity Approximately Gaussian



## Rapidity Distribution in the Thermal Model



## Superposition of Fireballs.

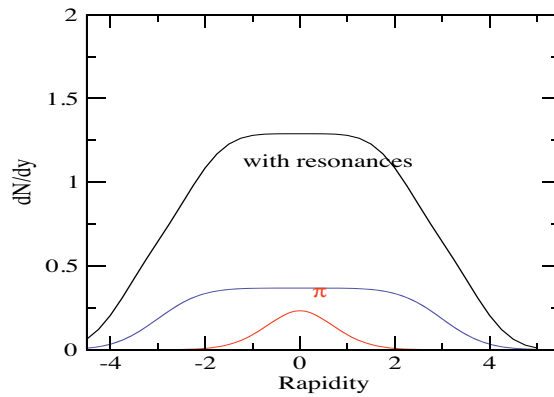
$$\frac{dN_i}{dy} = \int_{-Y}^Y dY_{FB} \rho(Y_{FB}) \frac{dN_i^0}{dy}(y - Y_{FB})$$

where

$$\frac{dN_i^0}{dy} = \frac{g_i V}{2\pi^2} \left[ \frac{2T^3}{\cosh^2(y - Y_{FB})} + \frac{2mT^2}{\cosh(y - Y_{FB})} + m^2 T \right] e^{\frac{\mu_i}{T}}$$

$$e^{-\frac{m}{T} \cosh(y - Y_{FB})}$$





## Superposition of Fireballs.

$$n_i = \int_{-\infty}^{\infty} dy \int_{-Y}^Y dY_{FB} \rho(Y_{FB}) \frac{dn_i^0}{dy}(y - Y_{FB})$$

$$n_i = \int_{-Y}^Y dY_{FB} \rho(Y_{FB}) \int_{-\infty}^{\infty} dy \frac{dn_i^0}{dy}(y - Y_{FB})$$

$$n_i = g_i V \frac{1}{2\pi^2} T m_i^2 K_2\left(\frac{m_i}{T}\right) e^{\frac{\mu_i}{T}} \int_{-Y}^Y dY_{FB} \rho(Y_{FB})$$

Equivalent to changing the volume  $V$ .



## Superposition of Fireballs.

$$\frac{n_i}{n_j} = \frac{\int_{-\infty}^{\infty} dy \int_{-Y}^Y dY_{FB} \rho(Y_{FB}) \frac{dn_i^0}{dy} (y - Y_{FB})}{\int_{-\infty}^{\infty} dy \int_{-Y}^Y dY_{FB} \rho(Y_{FB}) \frac{dn_j^0}{dy} (y - Y_{FB})}$$

$$\frac{n_i}{n_j} = \frac{n_i^0 \int_{-Y}^Y dY_{FB} \rho(Y_{FB})}{n_j^0 \int_{-Y}^Y dY_{FB} \rho(Y_{FB})}$$

$$\frac{n_i}{n_j} = \frac{n_i^0}{n_j^0} = \frac{m_i^2 K_2(m_i/T)}{m_j^2 K_2(m_j/T)} e^{(\mu_i - \mu_j)/T}$$

Effects Cancel Out in Ratios.

