

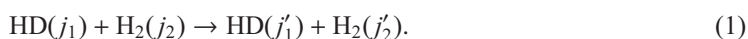
## HD + *ortho*-/*para*-H<sub>2</sub> rotational energy transfer non-reactive scattering at low astrophysical temperatures: 2 K ≲ T ≲ 300 K

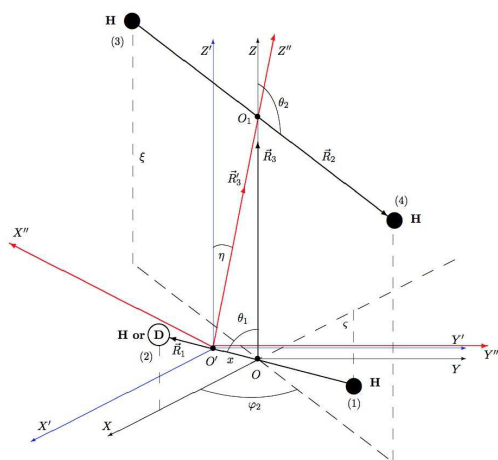
Renat A. Sultanov<sup>1,a</sup> and Dennis Guster<sup>1,b</sup><sup>1</sup>Department of Information Systems, BCRL & Integrated Science and Engineering Laboratory Facility (ISELF) at St. Cloud State University, St. Cloud, MN, USA

**Abstract.** State-resolved integral cross sections and corresponding thermal rate coefficients are computed for the HD + *ortho*-/*para*-H<sub>2</sub> rotational energy transfer collision at low temperatures of astrophysical interest: 2 K ≲ T ≲ 300 K. A recent original full-dimensional H<sub>2</sub>-H<sub>2</sub> potential energy surface (PES) was adopted and appropriately modified for the current non-symmetrical four-atomic system.

### 1 Introduction

One of the most important few-body systems in the field of atomic and molecular physics is, probably, the non-reactive scattering problem between two hydrogen molecules, i.e. H<sub>2</sub>+H<sub>2</sub>. The four-atomic symmetrical system, H<sub>2</sub>+H<sub>2</sub>, represents the two simplest and lightest identical diatomic molecules. In astrophysical applications H<sub>2</sub> can play a significant role, since it is the most abundant chemical species in the universe's molecular clouds. Consequently, the H<sub>2</sub>-H<sub>2</sub> interaction and collisions have always attracted significant theoretical and experimental attention. The system has been considered in many previous works, see for example [1–5]. After H<sub>2</sub>, the second most abundant primordial molecule is HD - deuterated hydrogen, where D is the hydrogen isotope <sup>2</sup>H - deuterium. Just like the H<sub>2</sub> + H<sub>2</sub> collision, the HD + H<sub>2</sub> energy transfer collision is also of significant importance in the astrophysics of the early Universe and in the modeling of pre-galactic clouds and planetary atmospheres. The HD molecule plays an important role in the chemistry of the interstellar medium [6] as well as the rotational and possibly vibrational energy transfer collisions between H<sub>2</sub> molecules. Further, collisions between H<sub>2</sub> and HD are important for the cooling of primordial gas and in the formation of first stars [7, 8]. In order to compute elastic/in-elastic scattering outputs, one needs to construct a PES between colliding molecules first. The HD-H<sub>2</sub> potential can be obtained from the initial H<sub>2</sub>-H<sub>2</sub> potential by shifting the original H<sub>2</sub> molecule's center of mass (c.m.) to the coordinate of the c.m. of the HD molecule. Once the exchange symmetry is broken in H<sub>2</sub>+H<sub>2</sub> by replacing the H by the D atom in one H<sub>2</sub> one has the new HD-H<sub>2</sub> PES. Thereby, one can derive the HD-H<sub>2</sub> PES from the original H<sub>2</sub>-H<sub>2</sub> PES employing all parts of the full HD-H<sub>2</sub> interaction including the HD's dipole moment. Using the HD-H<sub>2</sub> PES derived from the H<sub>2</sub>-H<sub>2</sub> PES of Hinde [4] we carry out full quantum calculations for collisions of rotationally excited HD and H<sub>2</sub> molecules, i.e. the process:

<sup>a</sup>e-mail: rasultanov@stcloudstate.edu<sup>b</sup>e-mail: dcguster@stcloudstate.edu



**Figure 1.** (Color online) Four-atomic system (1, 2, 3, 4) or H-D-H-H is shown together with its few-body Jacobi coordinates  $\{\vec{R}_1, \vec{R}_2, \vec{R}_3\}$ . The original Cartesian coordinate system is  $OXYZ$ . The center of mass of the original  $H_2$  molecule lies at point  $O$ .  $O'X'Y'Z'$  is a system which was shifted in parallel from the original system,  $O'$  lies in the center of mass of the actual HD molecule. The close-coupling equations are solved using the space-fixed coordinate system  $O'X''Y''Z''$ , which is obtained by rotation over the axis  $OY'$ . The vector  $\vec{R}_3$  connects the center of masses of the HD and  $H_2$  molecules. HD is the first molecule with a rotational constant  $B_e(1)=44.7 \text{ cm}^{-1}$  and quantum angular momentum of  $j_1$ ,  $H_2$  is the second molecule with rotational constant of  $B_e(2)=60.8 \text{ cm}^{-1}$  and uses momentum  $j_2$  in the system;  $R_1 = 0.7631 \text{ \AA}$  and  $R_2 = 0.7668 \text{ \AA}$  are fixed interatomic distances in each hydrogen molecule HD and  $H_2$  respectively.

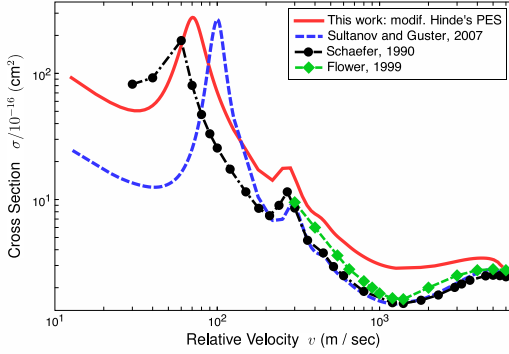
The scattering cross sections and their corresponding thermal rate coefficients are computed using a non-reactive quantum-mechanical close-coupling approach.

## 2 Quantum-mechanical method and results

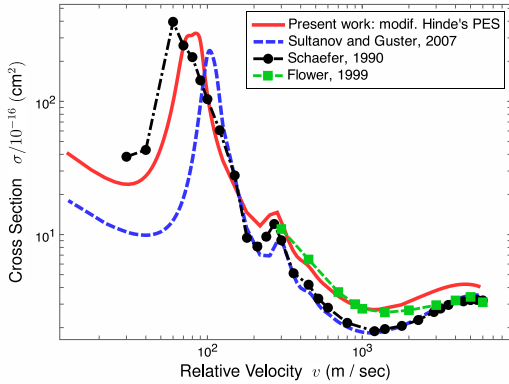
*1. Dynamical equations.* The details of the quantum-mechanical approach used in the current work has been described in few previous papers. See, for example, preprint [9] and references therein. Briefly, a four-atomic Schrödinger equation is numerically solved in the framework of the quantum-mechanical close-coupling approach [2]. The HD and  $H_2$  molecules are treated as linear rigid rotors. In our calculations, with the use of two PESs from works [3, 4] the bond length was fixed at 1.449 a.u. or 0.7668  $\text{Å}$  for the  $H_2$  molecule and 1.442 a.u. for HD which is 0.7631  $\text{Å}$ . The Schrödinger equation for the (12) + (34) collision in the c.m. frame, where (12) and (34) are diatomic molecules is:

$$\left[ \frac{\hat{P}_{\vec{R}_3}^2}{2\mathcal{M}_{12}} + \frac{\hat{L}_{\vec{R}_1}^2}{2\mu_1 R_1^2} + \frac{\hat{L}_{\vec{R}_2}^2}{2\mu_2 R_2^2} + V(R_1, \hat{R}_1, R_2, \hat{R}_2, \vec{R}_3) - E \right] \Psi(\hat{R}_1, \hat{R}_2, \vec{R}_3) = 0. \quad (2)$$

Here  $\hat{P}_{\vec{R}_3}$  is the momentum operator of the kinetic energy of the collision,  $\vec{R}_3$  is the collision coordinate, whereas  $\vec{R}_1$  and  $\vec{R}_2$  are relative vectors between atoms in the two diatomic molecules, and  $\hat{L}_{\vec{R}_{1(2)}}$  are the quantum-mechanical rotation operators of the rigid rotors,  $\mu_1$  and  $\mu_2$  are the reduced masses of the HD and  $H_2$  molecules and  $\mathcal{M}_{12}$  is the reduced mass of the two molecules. Also,  $V(R_1, \hat{R}_1, R_2, \hat{R}_2, \vec{R}_3)$  is the PES of the four-atomic system (1234),  $E$  is the total energy in the c.m. system, and  $\hat{R}_{1(2)}$  are the angles of orientation for rotors (12) and (34), respectively, it is shown in Fig. 1. The use and modification of the original  $H_2$ - $H_2$  PESs [3, 4],  $V(\vec{R}_1, \vec{R}_2, \vec{R}_3)$ , is discussed below and in greater detail in [9, 10]. The cross sections for excitation from  $HD(j_1, m_1) + H_2(j_2, m_2)$  to  $HD(j'_1, m'_1) + H_2(j'_2, m'_2)$  are summed over the final angular momentum projections ( $m'_1 m'_2$ ) and averaged over the initial projections ( $m_1 m_2$ ) of the HD and  $H_2$  molecules of angular momenta  $j_1$  and  $j_2$ :  $\sigma(j'_1, j'_2; j_1 j_2, \varepsilon) = \pi / [(2j_1 + 1)(2j_2 + 1)k_{\alpha\alpha'}] \sum_{J j_{12} j'_{12} LL'} (2J + 1) |\delta_{\alpha\alpha'} - S_{\alpha\alpha'}^J(E)|^2$ . The kinetic energy is  $\varepsilon = E - B_1 j_1(j_1 + 1) - B_2 j_2(j_2 + 1)$ , where  $B_1 = 44.7 \text{ cm}^{-1}$  and  $B_2 = 60.8 \text{ cm}^{-1}$  are the rotational



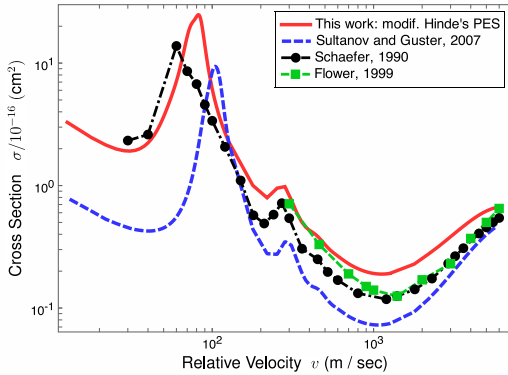
**Figure 2.** (Color online) The total cross section of the  $\text{HD}(1) + \text{H}_2(0) \rightarrow \text{HD}(0) + \text{H}_2(0)$  inelastic rotational energy transfer collision. The numbers in parentheses are the rotational quantum numbers of the two-atomic molecules.



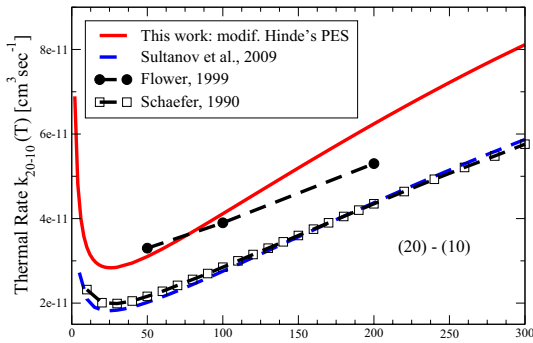
**Figure 3.** (Color online) The total cross section of the  $\text{HD}(2) + \text{H}_2(0) \rightarrow \text{HD}(1) + \text{H}_2(0)$  inelastic rotational energy transfer collision. The numbers in parentheses are the rotational quantum numbers of the two-atomic molecules.

constants of rigid rotors (12) and (34) respectively. Further,  $J$  is total angular momenta of the four-atomic system,  $\alpha \equiv (j_1 j_2 j_{12} L)$ , where  $j_1 + j_2 = j_{12}$  and  $j_{12} + L = J$ ,  $k_{\alpha\alpha'} = \sqrt{2M_{12}(E + E_\alpha - E_{\alpha'})}$  is the channel wavenumber and  $E_{\alpha(\alpha')}$  are rotational channel energies.

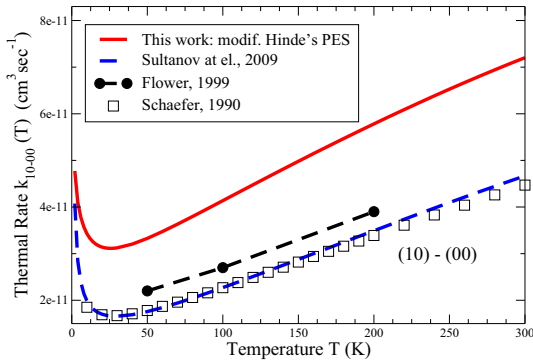
**2. Interaction potentials.** Based on the Born-Oppenheimer idea [11] any quantum-mechanical, semi-classical or, say, quasi-classical trajectory treatment of the  $\text{HD}+\text{H}_2$ ,  $\text{D}_2+\text{HD}$  or, say,  $\text{D}_2+\text{D}_2$  four-atomic systems should begin with consideration of the original  $\text{H}_2\text{-H}_2$  PES. The full six-dimensional function comprises a symmetrical surface over the  $OZ$  coordinate axis. In spherical coordinates the surface can be described by six variables:  $R_1, \theta_1, R_2, \theta_2, R_3$ , and  $\varphi_2$  [4]. The variables are shown in Fig. 1. In work [3] an  $\text{H}_2\text{-H}_2$  potential was formulated in a Cartesian coordinate system. In both works the  $\text{H}_2\text{-H}_2$  potential was derived in the framework of the Born-Oppenheimer model [11] and can basically be considered as a symmetrical interaction field. When considering non-reactive scattering problems with participation of hydrogen molecules, one needs to solve the Schrödinger Eq. (2) with the  $\text{H}_2\text{-H}_2$  potential  $V(\vec{R}_1, \vec{R}_2, \vec{R}_3)$ . The solution/propagation runs over the  $\vec{R}_3$  Jacobi vector, which can be directed over the  $OZ$  axis, this is shown in Fig. 1. Therefore, in the case of the symmetrical  $\text{H}_2+\text{H}_2$  and  $\text{D}_2+\text{D}_2$  collisions, one can use the original  $\text{H}_2\text{-H}_2$  PES as it is, i.e. without transformations. However, in the case of the non-symmetrical (symmetry-broken)  $\text{HD}+\text{H}_2/\text{D}_2$  or  $\text{HD}+\text{HD}$  scattering systems, one should also apply the original  $\text{H}_2\text{-H}_2$  interaction field or PES, but in this case the propagation (solution) of the Schrödinger equation should run over the corrected Jacobi vector  $\vec{R}_3'$  which is directed over the new  $O'Z''$  axis. This is shown in Fig. 1. In the current work, and in the preprint [9], we used two modified potentials from the papers [3, 4].



**Figure 4.** (Color online) The total cross section of the  $\text{HD}(2) + \text{H}_2(0) \rightarrow \text{HD}(0) + \text{H}_2(0)$  inelastic rotational energy transfer collision. The numbers in parentheses are the rotational quantum numbers of the two-atomic molecules.



**Figure 5.** (Color online) The rotational de-excitation thermal rate coefficient  $k_{ij \rightarrow i'j'}(T)$  for the following collision:  $\text{HD}(2) + \text{H}_2(0) \rightarrow \text{HD}(1) + \text{H}_2(0)$ .



**Figure 6.** (Color online) The rotational de-excitation thermal rate coefficient  $k_{ij \rightarrow i'j'}(T)$  for the following collision:  $\text{HD}(1) + \text{H}_2(0) \rightarrow \text{HD}(0) + \text{H}_2(0)$ .

**3. Results.** In order to solve the Schrödinger Eq. (2) the MOLSCAT program [12] has been used in our calculations. Because the  $\text{HD}+\text{H}_2$  total rotational energy transfer cross sections have shape resonances at the low energy sector, a large number of energy points were needed. We used up to 250 energy points in each computation for each specific rotational transition in the HD and  $\text{H}_2$  molecules considered. Convergence was obtained for the integral cross sections,  $\sigma(j'_1, j'_2; j_1, j_2, \epsilon)$ , with respect to the variation of the variables utilized in all considered collisions at different energies. For the intermolecular distance  $R_3$ , we used from  $R_{3min} = 3.0$  a.u. to  $R_{3max} = 30.0$  a.u. Our new rotational energy transfer integral cross sections are shown only for three different quantum-state transitions in the HD and  $\text{H}_2$  molecules. Specifically,  $\text{HD}(1) + \text{H}_2(0) \rightarrow \text{HD}(0) + \text{H}_2(0)$ ,  $\text{HD}(2) + \text{H}_2(0) \rightarrow \text{HD}(1) + \text{H}_2(0)$  and  $\text{HD}(2) + \text{H}_2(0) \rightarrow \text{HD}(0) + \text{H}_2(0)$  are presented in figures 2, 3 and 4. However, in

the recent preprint [9] more results for different rotational transitions in the HD and H<sub>2</sub> molecules are available together with the important thermal rate coefficients at different temperatures:  $k_{j_1 j_2 \rightarrow j'_1 j'_2}(T)$  [9]. The shape resonances in the low velocity region are reproduced, which are important in the cooling of the astrophysical media [6, 7]. Two thermal rate coefficients  $k_{j_1 j_2 \rightarrow j'_1 j'_2}(T)$  [9] are presented in figures 5 and 6 as examples. Again, one can find more results in [9]. The new results are compared with the older quantum-dynamical calculations of Schaefer [7] and Flower [8]. The computational results from the older work [10] are also shown. These data have been obtained by using a PES obtained from a modification of the BMKP H<sub>2</sub>-H<sub>2</sub> PES [3]. One can see that the new cross sections obtained with the Hinde's PES have the same structure and shape, but also have substantially larger ( $\sim 60\%$ ) values at medium energies ( $v \gtrsim 200$  m/sec) when compared with the results obtained from the modified BMKP PES [3] and the older result from Schaefer's calculations [7]. The cross sections from Ref. [8] are available only for higher collision velocities, i.e.  $v \geq 300$  m/s. In summary, state-resolved close-coupling quantum-mechanical calculations for rotational de-excitation cross-sections and corresponding thermal rate coefficients of the HD+*o*-/*p*-H<sub>2</sub> collisions are computed using a linear rigid rotor model for the HD and H<sub>2</sub> molecules. The symmetrical H<sub>2</sub>-H<sub>2</sub> PES of Ref. [4] has been appropriately adopted for the current non-symmetrical HD+H<sub>2</sub> system by appropriate translation and rotation. These geometrical operations lead to a new set of angle variables  $\theta'_1, \theta'_2$  and  $\varphi'_2$  for the Jacobi few-body coordinates, a new length of the intermolecular distance  $\vec{R}'_3$  and, as a result, to a new HD-H<sub>2</sub> PES [9]. The new results can be used for subsequent consideration (at different astrophysical conditions) the astrophysical HD-cooling function,  $\Lambda_{\text{HD}}(T)$ , which is important in the astrophysics of the early universe [6]. Also, the substantially larger cross sections and rates in HD+*o*-/*p*-H<sub>2</sub> found in the current work and in [9] with the use of the modified Hinde PES are consistent with previous results for H<sub>2</sub>+H<sub>2</sub> [13]. It has also been shown [13, 14] that full-dimensional dynamics which employ the Hinde PES [4] generally provide fairly good agreement with experiments for the low-lying vibrational states of H<sub>2</sub>. Therefore, it can be recommended that the new results presented in [13, 14] for H<sub>2</sub>+H<sub>2</sub>, and in [9] for HD+H<sub>2</sub>, which employ the Hinde PES, can be adopted in the astrophysical models.

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