

# Direct virtual photon production in Au+Au collisions with STAR BES-II data

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**Abstract.** In relativistic heavy-ion collisions, electromagnetic probes such as direct photons have the advantage of escaping from the emission sources without suffering strong interactions. Consequently, these photons, which are not from hadron decay, carry the information about the properties and dynamics of the hot and dense medium created in heavy-ion collisions. We report the measurements of the direct virtual photons from STAR experiment, based on the data collected in the Beam Energy Scan Phase II (BES-II) Program at RHIC. The  $p_T$ -integrated yields exhibit strong dependence on multiplicity and scale with  $(dN_{ch}/d\eta)^\alpha$ , where  $\alpha = 1.43 \pm 0.04$  (stat.)  $\pm 0.04$  (sys.). The  $p_T$ -dependent virtual photon mass spectra in  $\sqrt{s_{NN}} = 54.4$  GeV Au+Au collisions are measured for the first time. The results reveal that the observed mass shape of virtual photons varies with  $p_T$ , indicating a transition in their production mechanism from an equilibrated system to a non-equilibrated one as  $p_T$  increases.

## 1 Introduction

Direct photons are electromagnetic probes produced during all stages of ultra-relativistic heavy-ion collisions, escaping from the generated hot and dense medium without experiencing strong interactions. They are prompt photons from initial hard parton scattering and thermal photons from the thermal radiation of the medium. Prompt photons dominate the direct photon yield at high  $p_T$ , providing a good way to understand partonic dynamics and testing perturbative QCD theory. At low  $p_T$ , where thermal photons dominate, direct photons serve as penetrating probes of quark-gluon plasma (QGP), probing the energy density, temperature, and collective motion of the hot and dense medium generated from the collisions [1].

In this contribution, the measurements of direct virtual photon are presented for the first time at  $\sqrt{s_{NN}} = 14.6$  and  $19.6$  GeV in different centrality classes from Au+Au collisions from the STAR experiment, together with the previous results at  $\sqrt{s_{NN}} = 27$  and  $54.4$  GeV. The direct virtual photon mass distributions in different  $p_T$  regions in  $\sqrt{s_{NN}} = 54.4$  GeV Au+Au collisions have been studied, aiming a more comprehensive understanding of direct photon production mechanism.

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## 2 Experiment and analysis

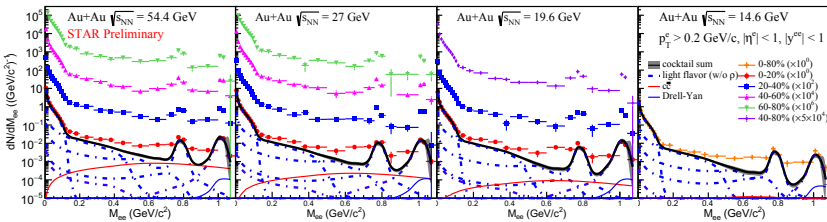
The presented analyzes use the data from Au+Au collisions at  $\sqrt{s_{NN}} = 14.6, 19.6, 27,$  and  $54.4$  GeV. Charged particle tracks and the average ionization energy loss of particles per unit length ( $\langle dE/dx \rangle$ ) are reconstructed by the Time Projection Chamber (TPC). Good electron identification has been achieved using both the TPC and Time-of-Flight (TOF) detectors. Both TPC and TOF cover the full azimuthal angle ( $-\pi < \phi < \pi$ ) and mid-pseudorapidity ( $|\eta| < 1$ ).

The dielectron invariant mass spectra are obtained by comparing randomly paired opposite-sign (+-) and like-sign (+/--) electrons/positrons. TPC tracking efficiency, TOF matching efficiency, and electron identification efficiency corrections were applied [3, 4]. Cocktail is the simulated dielectron background that originates from final-state hadron decays, including two-body and Dalitz decay of  $\pi^0$ ,  $\eta$ ,  $\eta'$ ,  $\omega$ ,  $\phi$ , and  $J/\psi$ ; semi-leptonic decay of charm; and Drell-Yan process. In particular, since  $\eta$  influence the direct photon yield largely and there is no  $\eta$  yield at these collision energies, the  $\eta/\pi^0$  ratio is constrained by worldwide data, fixed at  $0.470 \pm 0.017$  for  $p_T = 5$  GeV/c.

## 3 Results

### 3.1 Dielectron invariant mass spectra

Figure 1 presents the dielectron invariant mass spectra and cocktail within the STAR acceptance in Au+Au collisions from the STAR BES-II program. A clear enhancement caused by direct virtual photons and in-medium  $\rho^0$  can be observed in this presented mass region in 0-20% centrality.



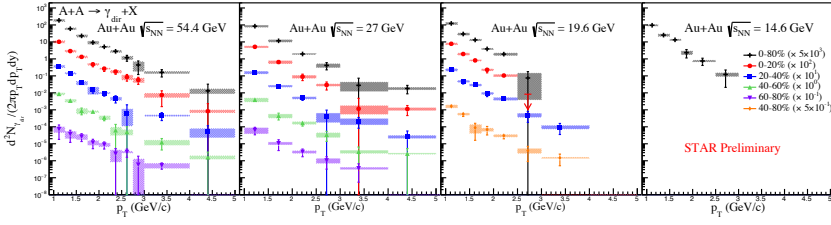
**Figure 1.** Dielectron invariant mass spectra and cocktail distributions within STAR acceptance from Au+Au collisions at  $\sqrt{s_{NN}} = 14.6, 19.6, 27, 54.4$  GeV.

### 3.2 Direct photon $p_T$ spectra and yields ( $dN/dy$ )

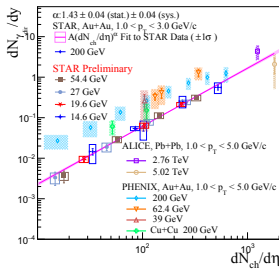
The direct virtual photon yield is extracted by internal method via a two-component fitting to the invariant mass distribution in different  $p_T$  regions,

$$\frac{dN_{ee}^2}{dM_{ee} dp_T} = r * f_{dir} + (1 - r) * f_{ckt}, \quad (1)$$

where  $r$  represents the fraction of the direct virtual photons to inclusive photons,  $f_{dir}$  and  $f_{ckt}$  represent the invariant mass distributions of direct virtual photons and the cocktail in different  $p_T$  regions, respectively. The direct virtual photon  $p_T$  spectra, shown in Fig. 2, are extracted by fitting the dielectron mass spectra in  $p_T$  intervals. The  $p_T$ -integrated yields from



**Figure 2.** Direct virtual photon  $p_T$  spectra in Au+Au collisions at  $\sqrt{s_{NN}} = 14.6, 19.6, 27, 54.4$  GeV in different centrality classes. Statistical and systematic uncertainties are shown as vertical lines and shaded boxes, respectively. The spectra are scaled for clarity.



**Figure 3.** Direct photon yields versus charged particle multiplicity in Au+Au collisions from BES-II program. Solid points represent 14.6, 19.6, 27, 54.4 GeV direct photon yields ( $1.0 < p_T < 3.0$  GeV/c) respectively, open points represent 2.76 and 5.02 GeV ( $1.0 < p_T < 5.0$  GeV/c) results from ALICE [6, 7] and 39, 62.4, 200 GeV results ( $1.0 < p_T < 5.0$  GeV/c) from PHENIX [8, 9], pink line corresponds to the fit to data points from STAR.

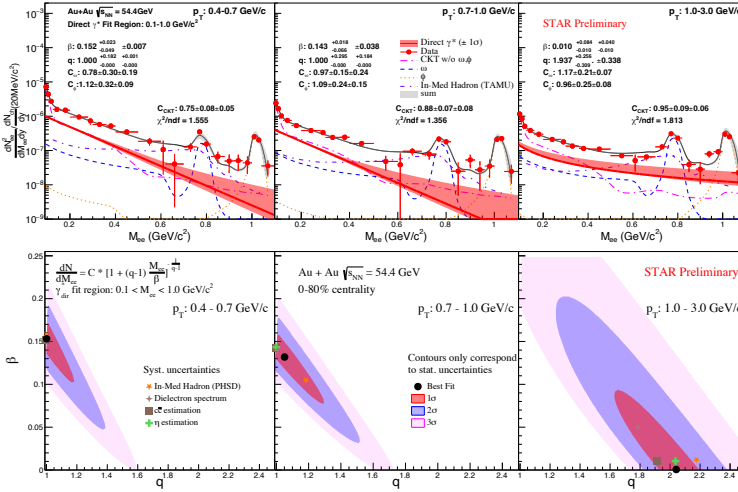
1 to 3 GeV/c of direct virtual photons are shown as a function of  $dN_{ch}/d\eta$  in Fig. 3, together with the results from  $\sqrt{s_{NN}} = 200$  GeV Au+Au collisions measured by STAR previously [5]. A strong multiplicity dependence has been observed for the direct virtual photon yields. An exponential function  $(dN_{ch}/d\eta)^\alpha$  is used to fit the direct virtual photon yields from STAR. The scaling power  $\alpha$  value is  $1.43 \pm 0.04(\text{stat.}) \pm 0.04(\text{sys.})$ .

### 3.3 Virtual photon mass distribution

To further study the virtual photon mass distribution, a function is shown as Eq. (2),

$$\frac{dN_{ee}}{dM_{ee}} = \gamma^{dir} + C_{ckt} * f_{ckt(w/o \omega, \phi)} + C_\omega * f_\omega + C_\phi * f_\phi + H_{in-medium}. \quad (2)$$

was used to fit the low mass dielectron continuum at  $\sqrt{s_{NN}} = 54.4$  GeV, where the direct virtual photon contribution ( $\gamma^{dir}$ ) is evaluated by  $C * f_{dir} = C * [1 + (q - 1) \frac{M_{ee}}{\beta}]^{-\frac{1}{q-1}}$ .  $f_{ckt(w/o \omega, \phi)}$ ,  $f_\omega$ ,  $f_\phi$  represents each component mass shape in the cocktail.  $C_{ckt}$ ,  $C_\omega$ ,  $C_\phi$  represents the amplitude.  $H_{in-medium}$  represents the dielectron contribution from the QGP medium estimated by TAMU. In this function, the parameter  $q$  determines the overall shape of the virtual photon mass distribution.  $f_{dir}$  approaches an exponential form when  $q$  is close 1, corresponding to a system close to thermal equilibrium. In contrast, when  $q$  is closer to 2,  $f_{dir}$  exhibits a power-law behavior, indicating a non-equilibrated system. The fitting results are presented in the upper panels of Fig. 4. As the transverse momentum increases, the extracted  $q$  increases from 1 to 2. This behavior suggests that the mass distribution of virtual photons changes from exponential to power-law as  $p_T$  increases, corresponding to the production from the later stage to the earlier stage. Furthermore, comparison between dielectron spectra and the fit in different  $p_T$  intervals, the contours of different confidence levels obtained from a  $\chi^2$ -minimization procedure by scanning different  $q$  and  $\beta$  are shown in the lower panels. The systematic uncertainty are shown in the lower panels of Fig. 4.



**Figure 4.** (Upper panels) Direct virtual photon mass distributions in Au+Au collisions at  $\sqrt{s_{NN}} = 54.4$  GeV in different  $p_T$  ranges, 0-80%. Red points represent the dielectron mass spectra in different  $p_T$  regions, red line is the direct virtual photon fit and red band represent  $1\sigma$  band. (Lower panels) The constraints on virtual photon distribution extracted by the comparison STAR measurement and fitting in different  $q$  and  $\beta$ .

## 4 Summary

Direct virtual photon  $p_T$  spectra have been presented for Au+Au collisions at  $\sqrt{s_{NN}} = 14.6, 19.6, 27,$  and  $54.4$  GeV measured with the STAR experiment. The yields ( $1 < p_T < 3$  GeV/c) exhibit a strong dependence on multiplicity and follow a common scaling with  $(dN_{ch}/d\eta)^\alpha$ , where  $\alpha = 1.43 \pm 0.04(\text{stat.}) \pm 0.04(\text{sys.})$ . The virtual photon mass spectra have been studied in detail. With  $p_T$  increasing from 0.4 to 3 GeV/c, the virtual photon mass distributions change from exponential to power-law, which hints at a production mechanism shifting from an equilibrated to a non-equilibrated system.

## Acknowledgments

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## References

- [1] E.V. Shuryak, Phys.Lett.B 78 (1978) 150.
- [2] G. David, Rept.Prog.Phys. 83 (2020) 4, 046301.
- [3] STAR Collaboration, arXiv:2402.01998 [nucl-ex].
- [4] J. Chen *et al.*, Nucl.Sci.Tech. 35 (2024) 12, 214.
- [5] L. Adamczyk, (STAR Collaboration), Phys.Lett.B 770 (2017) 451-458
- [6] B. Sahlmüller, (ALICE Collaboration), Nucl.Phys.A 956 (2016) 421-424.
- [7] D. Peresunko, (ALICE Collaboration), KnE Energ.Phys. 3 (2018) 210-216.
- [8] N. J. Abdulameer *et al.* (PHENIX Collaboration), Phys.Rev.C 107 (2023) 2, 024914.
- [9] N. J. Abdulameer *et al.* (PHENIX Collaboration), Phys.Rev.C 109 (2024) 4, 044912.